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Heat transfer analysis of unsteady MHD slip flow of ternary hybrid Casson fluid through nonlinear stretching disk embedded in a porous medium

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ABSTRACT

The original article's purpose is to assess transfer of heat exploration for unsteady magneto hydrodynamic slip flow of ternary hybrid Casson fluid via a nonlinear flexible disk placed within a perforated medium of a magnetic field in the presence. Unsteady nonlinearly stretched disk inside porous material causes flow to occur. In the investigation, convective circumstances on wall temperature are also considered. The governing equations (PDEs) are transformed into ordinary differential equations (ODEs) using appropriate transformations, and the Keller-box technique is employed for their solution. In forced convection, the variable radiation has no direct impact on fluid velocity, but it is noticed that in the case of aiding flow, fluid velocity rises with an increase in radiation parameter, and the opposite is true in the case of opposing flow. Furthermore, it is experiential that fluid concentration and velocity goes up in creative chemical reactions, and both profiles decrease in detrimental chemical reactions. Moreover, a slightly greater unsteadiness characteristic lowers fluid, concentration, temperature and velocity. Physical parameters' effects on fluid temperature, concentration, and velocity, as well as on wall shear stress, energy, and mass transfer rates, are studied.

1. Introduction

Normal heat transfer fluids cannot be used to meet cooling rate requirements. Such fluids are weaker thermal conductivities are the cause of this. The heat conductivity of ethylene glycol, water, and motor oil is poor. Due to high range of nanofluids they are used in mechanical engineering, industrial engineering and biomedical engineering. Cancer therapy, Cooling of electronic devices, and water distillation are due to nanofluids. Magneto nanofluids are used in magneto hydrodynamics, cancer therapy, cancer tumor treatment, power generators, and many others. The materials known as nanofluids are those that contain trace amounts of 1–100 nm-sized nanoparticles, also referred to as nanoparticles. The carbides, metals, and oxides, such as copper and gold, as well as oxides such as Titania, alumina and copper oxide, make up the nanoparticles. Nanofluids have also been used with diamond and carbon nanotubes. These particles can enhance the base liquids' thermo physical properties. Moreover, magnetic nanofluids are a solid structure with both liquid and magnetic properties. When dealing with conditions like cooling nuclear reactors with liquid sodium, magnetic field interaction with nanofluids is beneficial. Firstly Choi [1], investigated the thermal properties of nanoparticles. Buongiorno [2] researched a model for base fluid thermal properties. A model of a nanomaterial's stagnation point flow toward an extended piece was computed by Mustafa et al. [3]. Turkyilmazoglu [4] discovers the precise solution for magneto hydro dynamics slip flow energy and mass transfer. A model for convective boundary-layer flow of nanoparticles examined by Nield et al. [5]. While Shamshuddin et al. [6] simulate the Von Karman whirling bioconvection

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Nomen	clature	(1
		Р
<i>u</i> , <i>v</i> , <i>w</i> :	velocity components	μ
β:	Casson fluid parameter	α
ν_{nf} :	<i>n_n</i> : nanofluid kinematic viscosity	
ρ_{nf} :	nanofluid density	(7
k_{nf} :	nanofluid thermal conductivity	σ
B_0 :	magnetic field strength	k
T_w :	wall temperature	E
<i>k</i> *:	mean absorption coefficient	K
Pr:	Prandtl number	δ
Bi:	Biot number	
<i>M</i> :	Magnetic field parameter	

Table 1

The base liquid's and	nanomaterial's thermop	hysical properties.

Material	Base fluids			
	SA	Al_2O_3	Си	Ag
$ ho(kg/m^3)$	989	3970	8933	10,500
K(W/mK)	0.613	40	401	429
$c_p(J/kgK)$	4175	765	385	235
$\beta imes 10^{-5} (K^{-1})$	0.99	0.85	1.67	1.89
Pr	6	-	-	-

nanofluid flow using the Adomian decomposition technique. [7,8]. Furthermore, in [9–11] the effect of radiative heat on nanofluid flows. They address a variety of topics, including nanofluid flow in a parallel channel with thermal buoyancy, stability analysis of a stretching/ shrinking sheet, and axisymmetric nanofluid flow with reactions in porous media. These research provide light on convective heat transfer augmentation, dual solutions, and the behavior of nanofluids in complicated flow settings, all of which have practical relevance. Animasaun et al. [12] explores the thermal analysis of unsteady waterbased ternary hybrid nanofluids on wedges, considering

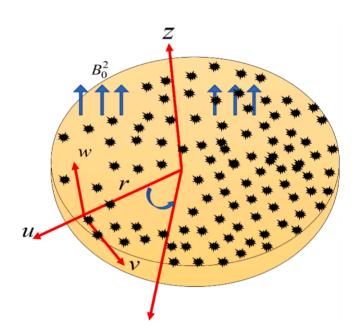


Fig. 1. Geometrical configuration of the model.

(r , z):	Cartesian coordinates
<i>P</i> :	pressure
μ_{nf} :	nanofluid dynamic viscosity
α_{nf} :	nanofluid thermal diffusivity
$(C_p)_{nf}$:	nanofluid specific heat transfer
T_{∞} :	ambient temperature
σ^* :	Stefan-Boltzmann constant
k_0 :	porosity parameter
Ec:	Eckert number
<i>K</i> :	Porous medium
δ:	slip parameter

bioconvection, wall stretching velocity, and different magnitudes of nanoparticle thermal conductivity. The dynamics of ternary-hybrid nanofluids on wedge surfaces under forced convection are examined in the paper by Xiu et al. [13] which takes into account variable water temperatures as well as small and large quantities of nanoparticles. (Table 1).

Hybrid nanofluids are classified as combinations or mixtures of two nanomaterials that are uniformly dispersed within a base liquid. When compared to nanofluids, hybrid nanofluids is more productive both monetarily and thermal. This new category of liquids denotes a liquid combination that has nano particles of two different solids distributed throughout one liquid. Hybrid nanoparticles have different thermal conductivity from base liquid. This kind of liquid has more uses related to the heating process and performs well in cooling where the temperature range is high. Hybrid nanofluids have numerous uses in the fields of refrigeration, solar energy, and industrial. Kshirsagar and Venkatesh [14] analysis of the setup of several apparatuses and investigation of the factors affecting hybrid nano-fluid performance. According to their conclusions, hybrid nanomaterials have higher heat conductivity than nanoparticles. A technique for calculating the conductivity of thermal of this fluids containing nanomaterials was scrutinized by Chen et al. [15]. The stability for dual solutions is examined in a simulation of timedependent radiative heat transport across a stretching/shrinking sheet of hybrid nanofluid by Tinker et al. [10]. The MHD flow of a viscous fluid across a flat plate under a low-pressure gradient is examined using the Homotopy Perturbation Method by Pattnaik, and Mishra [16]. In a computational research utilising a non-Fourier heat flux model, Shamshuddin et al. [17] examine heat transfer and viscous flow in a dual rotating extensible disc system. A numerical investigation on the impact of chemical reaction on magnetohydrodynamic flow across an exponential stretching sheet is done by Pattnaik et al. [18]. Due to the widespread usage of hybrid nanofluids, researchers have created a number of models for hybrid nanomaterials and have determined that they have a far higher thermal conductivity than nanofluids [19,20].

Casson fluid, a very effective non-Newtonian fluid, is used to assess inks, emulsions, food items, certain viscoelastic materials, and water colours. This fluid is one of the finest ones used today to mimic shear thinning liquids with rod-like components. In contrast to other fluid models like the power law, the Casson fluid model is capable of accepting complex rheological representative of a fluid. Cassion and Dash compute the fluid flow properties of a Casson fluid in a cylinder filled with a uniform porous medium [21,22]. Pham et al [23] mathematical study of the Cassion fluid led to the discovery of exit flows characterized by sudden contraction and a yield stress. Using velocity and temperature slip consequences, Amanullah et al. [24] investigated at the two-dimensional, laminar, energy and mass transfer of spontaneously convective micro liquid flow across a semi-infinite vertical plate surface. In a study to examine the effects of fluid viscosity, variable thermal conductivity, and nonlinear radiation on magnetohydrodynamic Casson nanofluid flow over a flat plate, Ghadikolaei et al. [25] found that velocity increases with an increase in variable thermal conductivity and viscosity while temperature and the volume fraction of nanoparticles decreases at a certain temperature. Das et al. [26] looked into the unsteady magnetohydrodynamics of chemically reactive double-diffusive Casson fluid on a level surface with heat and mass transfer. Mobile microorganisms can be beneficial for the suspension of nanofluids since they aid to preserve the distributed nanoparticles. It is important to note that the aforementioned study did not account for the lack of moving microorganisms.

On the basis of comprehensive literature review this article's goal aimed to evaluate heat transfer research into the porous medium of a magnetic field and the unsteady magneto hydrodynamic slip flow of a ternary hybrid Casson fluid through a nonlinear flexible disk. Flow is brought about by an unstable, nonlinearly stretched disk deep inside the porous material. Convective conditions on wall temperature are also taken into account in the study.

2. Mathematical formulation

We investigate a viscoelastic radiative fluid flow of a Casson trihybrid nanofluid, which consists of a mixture of nanoparticles, around a rotating disc. In this study, we consider the presence of three different types of nanoparticles within a single fluid. Additionally, we account for the influence of non-linear thermal radiation. The rotating disc is assumed to be in the xy-plane at z = 0, as shown in Figure 1. Our primary focus is to analyze the impact of the newly developed nanofluid model, known as "trihybrid nanofluid," on the fluid flow over the rotating disc. The continuity equation for the mixed convective flow of the Casson fluid is as follows:

$$\frac{\partial u}{\partial r} + \frac{\partial w}{\partial z} + \frac{u}{r} = 0 \tag{1}$$

$$\frac{\partial u}{\partial t} + \frac{u}{\partial r} \frac{\partial u}{\partial r} + \frac{w}{\partial z} \frac{\partial u}{r} - \frac{v^2}{r} = -\frac{\partial p}{\partial r} \frac{1}{\rho_{nf}} + \nu_{nf} \left(r^{-1} \frac{\partial u}{\partial r} + \frac{\partial^2 u}{\partial r^2} + \frac{\partial^2 u}{\partial z^2} - \frac{u}{r^2} \right) - \frac{B_0^2 \sigma_{nf} u}{\rho_{nf}}$$
(2)

$$\frac{\partial v}{\partial t} + \frac{u}{\partial r} + \frac{w}{\partial z} + \frac{uv}{r} = \nu_{nf} \left(\frac{\partial}{\partial r} \left(v \frac{1}{r} \right) + \frac{\partial^2 v}{\partial r^2} + \frac{\partial^2 v}{\partial z^2} \right) - \frac{B_0^2 \sigma_{nf} v}{\rho_{nf}} - \frac{\nu_{nf}}{k_0} v \tag{3}$$

$$\frac{\partial w}{\partial t} + \frac{w \partial w}{\partial z} + \frac{u \partial w}{\partial r} + = -\frac{\partial p}{\partial r} \frac{1}{\rho_{nf}} + \nu_{nf} \left(r^{-1} \frac{\partial w}{\partial r} + \frac{\partial^2 w}{\partial r^2} + \frac{\partial^2 w}{\partial z^2} \right)$$
(4)

$$\frac{\partial T}{\partial t} + w \frac{\partial T}{\partial z} + u \frac{\partial T}{\partial r} = \alpha_{nf} \left(\frac{\partial^2 T}{\partial r^2} + \frac{\partial^2 T}{\partial z^2} + \frac{1}{r} \frac{\partial T}{\partial r} \right) - \frac{1}{\left(\rho C_p\right)_{nf}} \partial q_r \frac{1}{\partial z}$$
(5)

By using the following BCs:

$$\left. \begin{array}{c} v = \Omega r, \ u = \xi r, \ w = 0, \\ T \rightarrow T_w, \ at \ z = 0 \\ v \rightarrow 0, \ u \rightarrow 0, \\ T \rightarrow T_\infty \ as \ z \rightarrow \infty, \end{array} \right\}$$
(6)

Here *P* for pressure, $\nu_{nf} = \frac{\mu_{nf}}{\rho_{nf}}$ kinematic viscosity and $\alpha_{nf} = \frac{k_{nf}}{(\rho C_p)_{nf}}$ represent thermal diffusivity.

The Rosseland diffusion approximation for radiation is:

$$q_r = \frac{-4\sigma^*}{3k^*} \frac{\partial T^4}{\partial z} \tag{7}$$

Stefan-Boltzmann constant is represent by σ^* . k^* for mean absorption coefficient. T_4 for temperature linear function. To expand T^4 using Taylor series with T^1 and remove higher terms, we can describe

$$T^4 \cong -3T^4_{\infty} + 4T^3_{\infty} T \tag{8}$$

Put Equation (9) and Equation (10) in Equation (4), as we get:

$$\frac{\partial T}{\partial t} + \frac{u\partial T}{\partial r} + \frac{w\partial T}{\partial z} = \alpha_{nf} \left(\frac{\partial^2 T}{\partial r^2} + r^{-1} \frac{\partial T}{\partial r} + \frac{\partial^2 T}{\partial z^2} \right) - \left(\frac{16\sigma^* T_{\infty}^3}{3k_1^* \left(\rho C_p\right)_{nf}} \right) \frac{\partial^2 T}{\partial z^2}$$
(9)

$$\frac{\partial T}{\partial t} + \frac{u\partial T}{\partial r} + \frac{v\partial T}{\partial z} = \frac{k_{thnf}}{\left(\rho c_p\right)_{thnf}} \left(\frac{\partial^2 T}{\partial r^2} + r^{-1}\frac{\partial T}{\partial r} + \frac{\partial^2 T}{\partial z^2}\right) - \frac{1}{\left(\rho c_p\right)_{thnf}} \frac{16\sigma^* T_{\infty}^3}{3k_1^*} \frac{\partial^2 T}{\partial z^2} + \frac{\nu_{thnf}}{\left(c_p\right)_{thnf}} \left(1 + \frac{1}{\beta}\right) \left(\left(\frac{\partial u}{\partial z}\right)^2 + \left(\frac{\partial v}{\partial z}\right)^2\right) + \frac{\sigma_{thnf} B_0^2}{\left(\rho c_p\right)_{thnf}} \left(u^2 + v^2\right) + \frac{Q(T - T_{\infty})}{\left(\rho c_p\right)_{thnf}},$$
(10)

by also provide the stream function (ψ) , described in terms of velocity in its standard notation, with a variable (η) , and the transformations shown below;

$$\eta = \sqrt{\frac{(n+1)\nu}{2\nu(1-\gamma t)}} r^{\frac{n-1}{2}} z, \ \psi = \sqrt{\frac{2c\nu}{(n+1)(1-\gamma t)}} r^{\frac{n+1}{2}} f(\eta), \ \vartheta = \frac{T-T_{\infty}}{T_f - T_{\infty}}$$
(11)

By using Eq. (8), the system of Eqs. (1–3) and (7)

$$A\left(\frac{\eta f^{'}}{n+1} + \frac{2f^{'}}{n+1}\right) = \left(1 + \frac{1}{\beta}\right)f^{''} + ff^{''} - \frac{2nf^2}{n+1} - (K+M)f^{'} + \lambda\vartheta$$
(12)

$$\left(1 + \frac{4}{3}R_d\right)\vartheta' + Prf\vartheta' - \frac{2(2n-1)}{n+1}Prf'\vartheta + Pr\left(1 + \frac{1}{\beta}\right)Ecf^{\prime 2} + PrMEcf^{\prime 2} + Pr\varepsilon\vartheta = PrA\left(\frac{2(2n-1)}{n+1}\vartheta + \frac{1}{n+1}\eta\vartheta'\right)$$
(13)

$$f'(0) = 1 + \left(1 + \frac{1}{\beta}\right) \delta f'(0),$$

$$\vartheta'(0) = -(1 - \vartheta(0))Bi$$

$$\vartheta(\infty) = 0, \ f'(\infty) = 0$$

Where
(14)

$$\begin{split} M &= \frac{2\sigma_f B_0^2}{\rho_f c(n+1)}, \ K = \frac{2v_f \varphi}{k_0(n+1)c}, \ \lambda = N_0 \sqrt{\frac{(n+1)c}{2v_f}}, \ Bi = \frac{h_0}{k_f} \left\{ \frac{2v_f}{c(n+1)} \right\}^{\frac{1}{2}}, \\ R_d &= \frac{4\sigma^* T_\infty^3}{\alpha k_1^*}, \ Pr = \frac{\mu_f (c_p)_f}{k_f}, \ Ec = \frac{u_w^2}{(c_p)_f (T_f - T_\infty)}, \ \varepsilon = \frac{Q_0}{(\rho c_p)_f (n+1)c}, \\ \delta &= N_0 \sqrt{\frac{(n+1)c}{2v_f}}, \\ \mu_{thom} &= \frac{\mu_f}{(1-\phi_{s1})^{2.5} (1-\phi_{s1})^{2.5} (1-\phi_{s1})^{2.5}} \end{split}$$

 $\rho_{thnf} = \rho_{f}(1 - \phi_{s2} - \phi_{s1} - \phi_{s3}) + \rho_{s1}\phi_{s1} + \rho_{s3}\phi_{s3} + \rho_{s2}\phi_{s2} \left(\rho C_{p}\right)_{thnf} = (1 - \phi_{s1} - \phi_{s2} - \phi_{s3})(\rho \beta_{T})_{bf} + (\rho \beta_{T})_{s2}\phi_{s2} + (\rho \beta_{T})_{s3}\phi_{s3} + (\rho \beta_{T})_{s1}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s2} + \rho_{s2}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s2}\phi_{s2} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} + \rho_{s2}\phi_{s1} + \rho_{s1}\phi_{s1} +$

$$\begin{split} \frac{K_{nf}}{K_{f}} = & \left\{ \frac{K_{s3} + 2K_{f} - 2\phi_{s3}\left(K_{f} - K_{s3}\right)}{K_{s3} + 2K_{f} + \phi_{s3}\left(K_{f} - K_{s3}\right)} \right\}, \frac{K_{hnf}}{K_{nf}} = \left\{ \frac{K_{s2} + 2K_{nf} - 2\phi_{s1}\left(K_{nf} - K_{s1}\right)}{K_{s1} + 2K_{nf} + \phi_{s1}\left(K_{nf} - K_{s1}\right)} \right\}, \\ \frac{K_{thnf}}{K_{hnf}} = \left\{ \frac{K_{s2} + 2K_{hnf} - 2\phi_{s2}\left(K_{hnf} - K_{s2}\right)}{K_{s2} + 2K_{hnf} + \phi_{s2}\left(K_{hnf} - K_{s2}\right)} \right\}. \end{split}$$

$$\frac{g_1}{g_2} \left(1 + \frac{1}{\beta} \right) f^{"} + f f^{'} - \frac{2n}{n+1} f^{'2} - \left(\frac{g_6}{g_2} M + \frac{g_1}{g_2} K \right) f^{'} + \lambda \frac{g_3}{g_2} \vartheta$$
$$= A \left(\frac{2}{n+1} f^{'} + \frac{1}{n+1} \eta f^{'} \right)$$
(16)

$$\begin{split} \sigma_{nf} &= \sigma_f \left\{ \frac{\sigma_{s3}(1+2\phi_{s3}) + \sigma_f(1-2\phi_{s3})}{\sigma_{s3}(1-\phi_{s3}) + \sigma_f(1+\phi_{s3})} \right\}, \sigma_{hnf} = \sigma_{bf} \left\{ \frac{\sigma_{s2}(1+2\phi_{s2}) + \sigma_{nf}(1-2\phi_{s2})}{\sigma_{s2}(1-\phi_{s2}) + \sigma_{nf}(1+\phi_{s2})} \right\},\\ \sigma_{thnf} &= \sigma_{hnf} \left\{ \frac{\sigma_{s1}(1+2\phi_{s1}) + 2\sigma_{hnf}(1-2\phi_{s1})}{\sigma_{s1}(1-\phi_{s1}) + \sigma_{hnf}(2+\phi_{s1})} \right\} \end{split}$$

$$g_1 = \frac{1}{(1 - \phi_{s1})^{2.5} (1 - \phi_{s2})^{2.5} (1 - \phi_{s3})^{2.5}}$$

$$g_{2} = 1 - \phi_{s2} - \phi_{s1} - \phi_{s3} + \phi_{s1} \frac{\rho_{s1}}{\rho_{f}} + \phi_{s3} \frac{\rho_{s3}}{\rho_{f}} + \phi_{s2} \frac{\rho_{s2}}{\rho_{f}} \\ g_{4} = 1 - \phi_{s1} - \phi_{s2} - \phi_{s1} - \phi_{s2} - \phi_{s3} + \phi_{s2} \frac{(\rho\beta_{T})_{s2}}{(\rho\beta_{T})_{f}} + \phi_{s3} \frac{(\rho\beta_{T})_{s3}}{(\rho\beta_{T})_{f}} \\ g_{4} = 1 - \phi_{s1} - \phi_{s2} + \phi_{s1} \frac{(\rhoC_{p})_{s1}}{(\rhoC_{p})_{bf}} - \phi_{s3} + \phi_{s2} \frac{(\rhoC_{p})_{s2}}{(\rhoC_{p})_{bf}} + \phi_{s3} \frac{(\rhoC_{p})_{s3}}{(\rhoC_{p})_{bf}} \\ g_{5} = 1 - \phi_{s1} - \phi_{s2} - \phi_{s3} + \phi_{s2} \frac{(\rhoC_{p})_{s2}}{(\rhoC_{p})_{bf}} + \phi_{s3} \frac{(\rhoC_{p})_{s3}}{(\rhoC_{p})_{bf}} + \phi_{s3} \frac{(\rhoC_{p})_{s3}}{(\rhoC_{p})_{s3}} + \phi_{s3} \frac{(\rhoC_{p})_{s3}}{(\rhoC_{p}$$

$$g_{5} = \frac{K_{thoff}}{K_{f}}g_{6} = \frac{\sigma_{thoff}}{\sigma_{f}}$$

$$\frac{\partial u}{\partial t} + u\frac{\partial u}{\partial x} + v\frac{\partial v}{\partial y} = \frac{g_{1}}{g_{2}}\left(1 + \frac{1}{\beta}\right)\frac{\partial^{2}u}{\partial y^{2}} - \left(\frac{g_{6}}{g_{2}}\frac{\sigma_{thoff}B_{0}^{2}}{\rho_{thoff}} + \frac{g_{1}}{g_{2}}\frac{\nu_{thoff}\varphi}{k_{1}}\right)u + \frac{g_{3}}{g_{2}}g(\beta_{T})_{thoff}(T - T_{\infty})$$
(15)

$$\frac{g_5}{g_4}\left(1+\frac{4}{3}R_d\right)\vartheta' + Prf\vartheta' - \frac{2(2n-1)}{n+1}Prf\vartheta + \frac{g_1}{g_4}Pr\left(1+\frac{1}{\beta}\right)Ecf^2 + \frac{g_6}{g_4}PrMEcf^2 + \frac{1}{g_4}Pr\varepsilon\vartheta = PrA\left(\frac{2(2n-1)}{n+1}\vartheta + \frac{1}{n+1}\eta\vartheta'\right)$$
(17)

3. Results and discussion

Through the use of a nonlinear stretch disk embedded in a porous medium, this work aims to enhance heat transfer analysis of unsteady magneto hydrodynamic slip flow of ternary hybrid Casson fluid. To

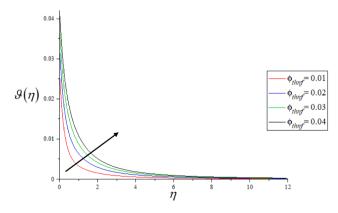


Fig. 2. Consequence of $\phi_{\textit{thnf}}$ on Temperature of fluid.

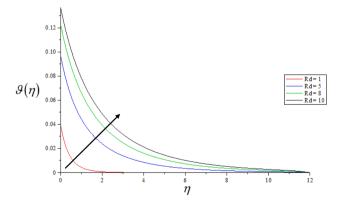
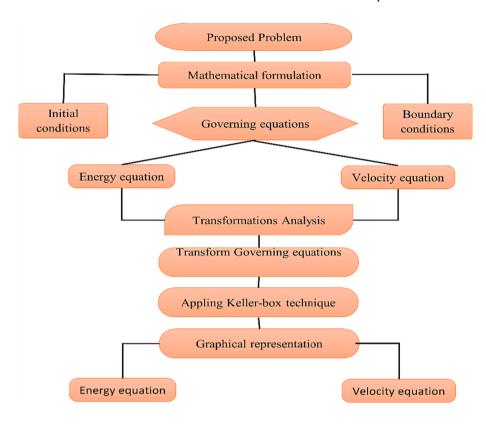


Fig. 3. Consequence of *Rd* on temperature of fluid.

resolve the problem, we used a Matlab program and the Keller-box strategies as the details are in the following flow chart. Flow chart. Figure 3 and figure 13 are plotted to displays the consequence of *Rd* on temperature and velocity profiles. The greater value of radiation parameter enhances the temperature profile and fluid velocity. Physi-



4. Temperature profile

The influences of several parameters are plotted for temperature profile in figures 2-11 and velocity profile in figures 12-21.

The volume fraction effect of tri-hybrid nanofluids is presents in figure 2 and figure 12 for temperature and velocity respectively. The enhancement of volume fractional enhance the temperature profile while decrease the velocity of the fluid. This result is due to the tri nanofluid being added to the base liquid, which strengthens the viscous forces in the trihybrid nanofluid.

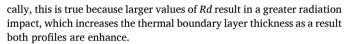


Figure 4 and figure 14 demonstrate the effect of n on both velocity and temperature profiles. The thickness of the thermal and velocity boundary layers are both rapidly reduced as the parameter n is increased.

A declining tendency in the velocity profile has been seen as a result of the Lorentz force, which is generated as M increases and slows the fluid's motion. The temperature profile of the fluid is also impacted by the Lorentz force. The thermal boundary layer thickens because electromagnetic forces are stronger than viscous forces at larger values of M. Figure 15 and Figure 5 show these outcomes, respectively.

The effects of K on the temperature and velocity profiles were highlighted in Figures 6 and 16, respectively. Physically, the size of the

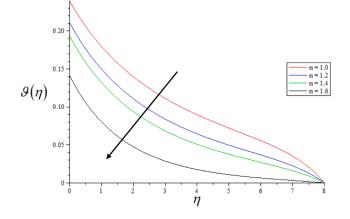


Fig. 4. Consequence of *n* on Temperature of fluid.

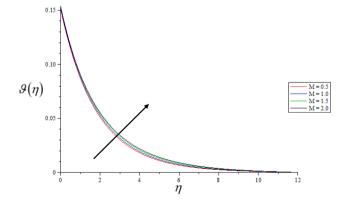


Fig. 5. Consequence of *M* on Temperature of fluid.

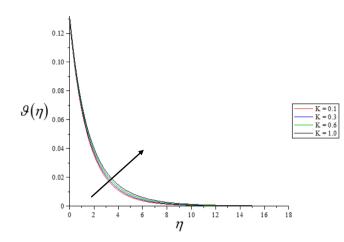


Fig. 6. Consequence of K on Temperature of fluid.

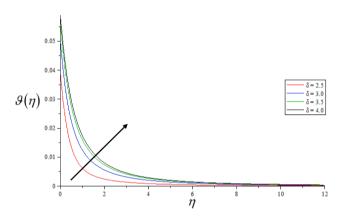


Fig. 7. Consequence of δ on Temperature of fluid.

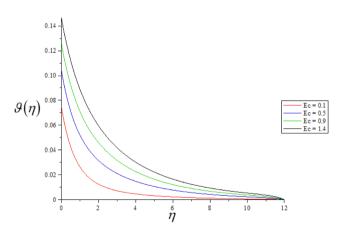


Fig. 8. Consequence of Ec on Temperature of fluid.

pores in porous media widens as K rises; in this scenario, the drag force acting against the flow direction causes the velocity boundary layer thickness to decrease and the velocity to be retarded. In contrast to unsteady flow, stable flow exhibits a stronger impact of K and a thicker thermal boundary layer.

Figure 7 and figure 17 shows the outcomes of δ on temperature and velocity profiles. It observed that the slipping parameter enhance the temperature of the fluid while the rapidly decrease the fluid velocity. This behavior can be explained by the fact that as flow near the disk decreases and more fluid is allowed to slip over the disk, the slip effect toward the free stream becomes less strong.

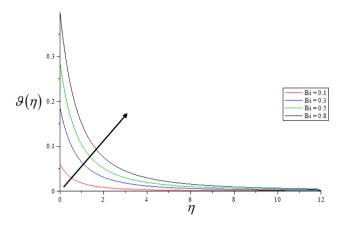


Fig. 9. Consequence of Bi on Temperature of fluid.

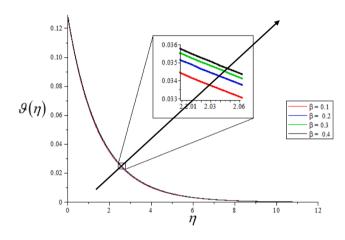


Fig. 10. Consequence of β on Temperature of fluid.

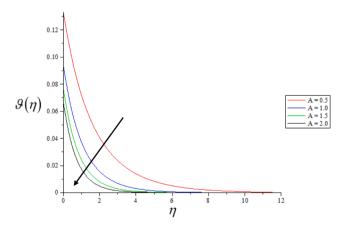


Fig. 11. Consequence of A on Temperature of fluid.

Figures 8 and 18 show how the Eckert number *Ec* responds to heat transfer and fluid velocity. As the amplitude of rises *Ec*, a rinsing tendency is visible in both profiles. Since it represents the proportion of the boundary layer's kinetic energy to its enthalpy difference, this is a physically sound statement. Since a positive signifies that heat has been transferred from the plate to the fluid, the temperature of the fluid rises as viscous heat dissipation increases.

Consequences of fluid temperature and velocity profile are shown in Figures 9 and 19. The ratio of the thermal resistance inside the disk to that of the boundary layer of the hot fluid at the surface's bottom is known as the Biot number. Temperature is an increasing function of *Bi* in

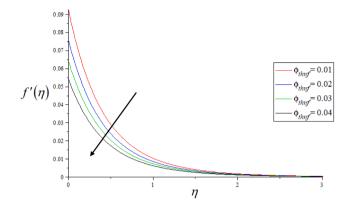


Fig. 12. Consequence of ϕ_{thnf} on velocity of fluid.

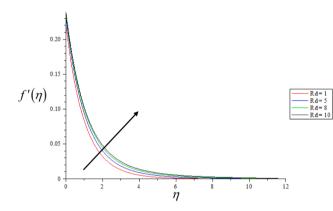


Fig. 13. Consequence of Rd on velocity of fluid.

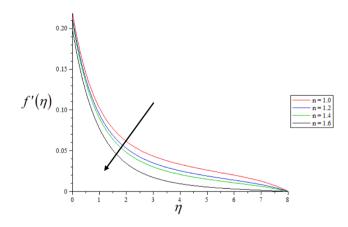


Fig. 14. Consequence of *n* on Velocity of fluid.

each of these scenarios. Convection at the surface is strengthened by the force, which causes a rise in surface temperature and, as a result, an increase in fluid velocity.

Clearly, Figure 10 and figure 20 depicts that the greater value of β caused the enhancement of heat transfer profile and fluid velocity close to the plate while fall the velocity gone from the plate. This occurrence demonstrates how a drop in yield stress can result in an rise in the thickness of the thermal boundary layer. Furthermore, the impact is exacerbated for linearly expanding disks that are not completely stable.

Lastly, figures 11 and 21 show the analysis of *A* on both temperature and velocity profiles. Both velocity and temperature fluid are

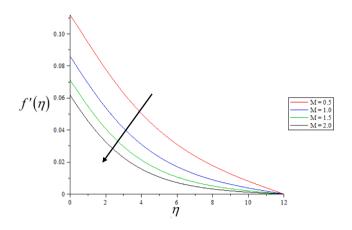


Fig. 15. Consequence r of M on Velocity of fluid.

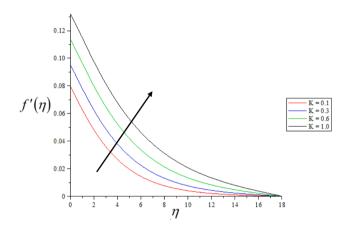


Fig. 16. Consequence of K on Velocity of fluid.

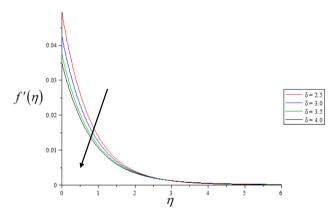


Fig. 17. Consequence of δ on Velocity of fluid.

appreciated to fall as *A* enhances. We also see that the boundary layer is thinner for greater *A* amplitudes.

The assessment of skin frictional of the present papers is compared with Nadeem et al. [27], Oyelakin et al. [28] and Ullah et al. [29] in Table 2, which shows a good agreement to existing literature. Table 3 present the comparison of heat transfer rate of the tri hybrid nanofluid to the classical and hybrid nanofluid, it is observed that the ternary hybrid nanofluid enhance the rate of heat transfer up-to 38.88529% as compared to the classical fluid.

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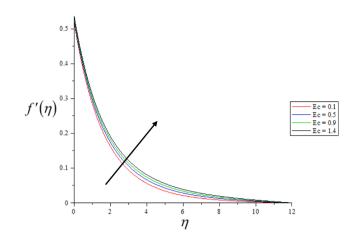


Fig. 18. Consequence of Ec on Velocity of fluid.

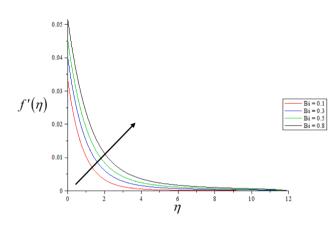


Fig. 19. Consequence of Bi on Velocity of fluid.

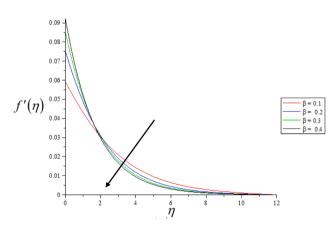


Fig. 20. Consequence of β on Velocity of fluid.

5. Conclusion

With convective conditions on wall temperature taken into consideration, the current work intends to analyse the heat transfer of unsteady magnetohydrodynamics slip flow of ternary hybrid Casson fluid through a nonlinear stretching disk embedded in a porous medium. The Kellerbox approach should be used to solve the issue. The core ideas of this book are as follows:

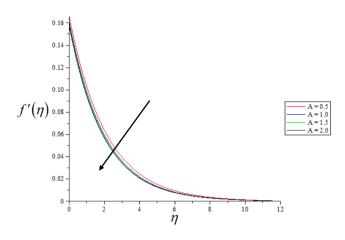


Fig. 21. Consequence of A on Velocity of fluid.

Table 2

Skin friction comparison for various values of M and β when $Bi \rightarrow \infty$, $n = 1, A = M = K = N = R_d = Sc = R = \lambda = \delta = \phi_1 = \phi_2 = 0$.

Μ	β	Nadeem et al. [27]	Oyelakin et al. [28]	Ullah et al. [29]	Present
0	5	-1.0954	-1.0944	-1.0955	-1.0955
0	1	-1.4142	-1.41421	-1.4144	-1.4144
10	5	-3.6331	-3.63318	-3.6332	-3.6332
10	1	-4.6904	-4.69042	-4.6904	-4.6904
100	5	-11.0091	-11.00909	-11.0091	-11.0091
100	1	-14.2127	-14.21267	-14.2127	-14.2127

Table 3

Nu Comparision of Tri Nano fluid with hybrid nano fluid, nanofluid and Classical fluid.

Name fluid	Nu	%age	
Classical fluid	1.543		
Nanofluid	1.743	12.96176	
Hybrid nanofluid	1.943	25.92353	
Tri nanofluid	2.143	38.88529	

- 1. By rising the values of *R*, then the result the velocity and concentration profile show the decreasing behavior.
- 2. Due to the radiation parameter, fluid velocity increases as the radiation parameter grows for a helping flow and drops as the radiation parameter increases for an opposing flow. The radiation parameter has no impact on fluid velocity in the presence of forced convection.
- 3. As a result of generative chemical reactions, fluid velocity and concentration rise, whereas destructive chemical reactions cause both characteristics to fall.
- 4. temperature and concentration are shown in the situation of unsteady flow with the effects of *M* and *K*.
- 5. Temperature is an increasing function of Bi1 in each of these situations. The reason for this is due to the fact that the asset of *Bi* rises the convective at the surface, which in turn leads to a rise in the temperature of the surface and so on the velocity of the fluid also enhance.

Declaration of Competing Interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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